

Chapter 10 Identical Particles

3. Fermi Dirac Distribution

In quantum statistics, Fermi–Dirac statistics describes distribution of particles in a system comprising many identical particles that obey the Pauli exclusion principle (The Pauli exclusion principle is the quantum mechanical principle that no two identical fermions (particles with half-integer spin) may occupy the same quantum state simultaneously.) Fermi–Dirac (F–D) statistics applies to identical particles with half odd integer spin in a system in thermal equilibrium. Additionally, the particles in this system are assumed to have negligible mutual interaction. This allows the many-particle system to be described in terms of singleparticle energy states. The result is the F–D distribution of particles over these states and includes the condition that no two particles can occupy the same state, which has a considerable effect on the properties of the system. Since F–D statistics applies to particles with half-integer spin,

these particles have come to be called fermions . It is most commonly applied to electrons, which

are fermions with spin $\frac{1}{2}$

No. of ways W in which n_i indistinguishable particles to place in g_i level with the condition that only one particle or no particle can be placed in each level i.e. identical particles that **obey the**

Pauli exclusion principle. (It is given $\sum_{i=1}^{i=l} n_i = N$ $\sum_{i=1}^{i=l} E_i n_i = U$.)

$$W = \sum_{i=1}^{l} \frac{|g_i|}{|n_i|g_i - n_i|}$$

Fermi-Dirac distribution of the particles among various states is given by

$$n_i = \frac{g_i}{\exp(\alpha + \beta E) + 1}$$

So Fermi Dirac distribution $f(E) = \frac{n_i}{g_i} = \frac{1}{e^{(\alpha+\beta E)} + 1} = \frac{1}{Ae^{\beta E} + 1}$

where
$$\beta = \frac{1}{k_B T}$$
 and $A = e^{\alpha} = \frac{V}{N} \left(\frac{2\pi m k_B T}{h^2}\right)^{3/2}$

when A.>> 1, Fermi-Dirac gas reduces to the Maxwell-Boltzmann gas. Fermi-Dirac gas is said to be weakly degenerate when A>1, degenerate when A<1 and strongly degenerate when A = 0.

Example: What is number of ways if two fermions have to adjust in energy state whose degeneracy is three.

Solution:
$$g_i = 3$$
, $n_i = 2$ $W = \frac{|g_i|}{|n_i|g_i - n_i|} = 3$

Two indistinguishable particles is shown by A,A

| Possible | 1 | 2 | 3 |
|-----------|---|---|---|
| selection | | | |
| 1 | A | A | 0 |
| 2 | 0 | Δ | Δ |
| Σ | 0 | | |
| 3 | А | 0 | А |
| | | | |

Fermions at High Temperature

A gas of low density and high temperature is known as a weakly degenerate gas (A > 1). In a Fermi gas, total number N of particles and total energy U are

 $g(E)dE = 2g\pi V \left(\frac{2m}{h^2}\right)^{3/2} E^{1/2} dE$ is density of state in three dimension with degeneracy

g = 2s + 1 where s is spin .

Total no of particle at temperature *T* is $N = \int_0^\infty f(E)g(E)dE = \int_0^\infty \frac{g(E)}{Ae^{\beta E} + 1}dE$

Total energy at temperature T $U = \int_0^\infty E f(E) g(E) dE = \int_0^\infty \frac{Eg(E)}{Ae^{\beta E} + 1} dE$

$$N = 2g\pi V \left(\frac{2m}{h^2}\right)^{3/2} \int_0^\infty \frac{E^{1/2}}{Ae^{\beta E} + 1} dE = 2g\pi V \left(\frac{2m}{h^2}\right)^{3/2} F_{3/2}(A)$$
$$U = 2g\pi V \left(\frac{2m}{h^2}\right)^{3/2} \int_0^\infty \frac{E^{3/2}}{Ae^{\beta E} + 1} dE = 2g\pi V \left(\frac{2m}{h^2}\right)^{3/2} F_{5/2}(A)$$

where $F_n(A) = \int_0^\infty \frac{E^{n-1}}{Ae^{\beta E} + 1} dE$

Let $\beta E = u$, then $\beta dE = du$ and we have $F_n(A) = \frac{1}{\beta_n} \int_0^\infty \frac{u^{n-1}}{Ae^u + 1} du = (k_B T)^n \Gamma(n) f_n(A)$

where $f_n(A) = \frac{1}{\Gamma(n)} \int_0^\infty \frac{u^n - 1}{Ae^u + 1} du$ and $\Gamma(n)$ is a gamma function.

$$N = 2g\pi V \left(\frac{2m}{h^2}\right)^{3/2} (k_B T)^{3/2} \Gamma(3/2) f_{3/2}(A)$$

$$= gV \left(\frac{2\pi m k_B T}{h^2}\right)^{3/2} f_{3/2}(A) = \frac{gV}{\lambda^3} f_{3/2}(A)$$

$$U = 2g\pi V \left(\frac{2m}{h^2}\right)^{3/2} (kT)^{3/2} \Gamma(5/2) f_{5/2}(A)$$

$$= \frac{3}{2} kTgV \left(\frac{2\pi m k_B T}{h^2}\right)^{3/2} f_{5/2}(A) = \frac{3}{2} k_B T \frac{gV}{\lambda^3} f_{5/2}(A)$$

where $f_n(A) = \sum_{r=1}^{\infty} (-1)^{r-1} \frac{1}{A^r} \frac{1}{r^n}$

Very weakly degenerate gas

A gas with very large value of A i.e. at very high temperature T is known as very weakly degenerate gas. For that $f_r(A) = A^{-1}$.

$$N = \frac{gV}{\lambda^3} f_{3/2}(A) = \frac{gV}{\lambda^3} \frac{1}{A}$$
$$U = \frac{3}{2} k_B T \frac{gV}{\lambda^3} f_{5/2}(A) = \frac{3}{2} k_B T \frac{gV}{\lambda^3} \frac{1}{A}$$

Thus, we have $\frac{U}{N} = \frac{3}{2}k_BT$ $U = \frac{3}{2}Nk_BT$

and the specific heat at constant volume c_v is $c_V = \left[\frac{\partial U}{\partial T}\right]_V = \frac{3}{2}Nk_B$ so at very high temperature

Fermi gas behave as ideal gas.

Weakly Degenerate Gas

For A>1, the gas is said to be weakly degenerate and equations and give

$$N = \frac{gV}{\lambda^3} f_{3/2}(A) \qquad \qquad U = \frac{3}{2} k_B T \frac{gV}{\lambda^3} f_{5/2}(A)$$

Therefore,

$$\begin{split} & \frac{U}{N} = \frac{3}{2} k_B T \frac{f_{5/2}(A)}{f_{3/2}(A)} \qquad U = \frac{3}{2} N k_B T \frac{f_{5/2}(A)}{f_{3/2}(A)} \\ & = \frac{3}{2} N k_B T \left[1 + \frac{1}{4\sqrt{2}} \frac{N}{V} \left(\frac{h^2}{2\pi m k_B T} \right)^{3/2} + \left(\frac{1}{16} - \frac{4}{9\sqrt{3}} \right) \frac{N^2}{V^2} \left(\frac{h^2}{2\pi m k_B T} \right)^{\frac{3}{2}} \right] \\ & c_V = \left[\frac{\partial U}{\partial T} \right]_V = \frac{3}{2} k_B N \left[1 - \frac{1}{8\sqrt{2}} \frac{N}{V} \left(\frac{h^2}{2\pi m k_B T} \right)^{3/2} - \left(\frac{1}{8} - \frac{4}{9\sqrt{3}} \right) \frac{N^2}{V^2} \left(\frac{h^2}{2\pi m k_B T} \right)^{3} \right] \end{split}$$

Fermions at Low Temperature

Strongly degenerate Fermi gas A < 1

The Fermi-Dirac energy distribution is

$$f(E) = \frac{1}{\exp(E-\mu)/k_B T + 1}$$

where μ is the chemical potential which is a function of T, i.e; $\mu = \mu(T)$. the gas is strongly degenerate (A = 0) at T = 0. at T = 0, where $\mu = \mu(0) = E_F$. The limiting chemical potential is known as the Fermi energy E_F of the gas and the distribution function can be written as

$$f(E) = \frac{1}{e^{(E-E_F)/k_BT} + 1}$$

Strongly degenerate Fermi gas at T = 0At T = 0, when E < E_F, we have $f(E) = \frac{1}{e^{-\infty} + 1} = \frac{1}{0 + 1} = 1$ At T = 0, when E > E_F, we have $f(E) = \frac{1}{e^{\infty} + 1} = \frac{1}{\infty + 1} = 0$ Fermi function f(E) versus E atT The number of energy states in the energy range from E to E + dE $g(E)dE = 2g\pi V \left(\frac{2m}{h^2}\right)^{3/2} E^{1/2} dE$ Here, g is the spin degeneracy, g = (2s+1), where s is the spin

quantum number of a particle. The number of particles in the energy range from E to E + dEat T = 0 is

$$n(E)dE = \begin{cases} f(E)g(E)dE = 2g\pi V \left(\frac{2m}{h^2}\right)^{3/2} E^{1/2}dE & \text{for } E < E_F \\ 0 & \text{for } E > E_F \end{cases}$$

$$N = \int_0^{E_F} n(E) dE = 2g\pi V \left(\frac{2m}{h^2}\right)^{3/2} \int_0^{E_F} E^{1/2} dE = \frac{4g\pi V}{3} \left(\frac{2m}{h^2}\right)^{3/2} E_F^{3/2}$$

Thus, the Fermi energy is $E_F(0) = \left(\frac{3N}{4\pi gV}\right)^{2/3} \frac{h^2}{2m}$ and the Fermi temperature T_F is defined as

$$T_F = \frac{E_F}{k_B}$$

The Fermi momentum p_F is given by

$$p_F = (2mE_F)^{1/2} = \left(\frac{3N}{4\pi gV}\right)^{1/3} h$$

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Total energy of the gas at T = 0 is

$$U = \int_0^\infty En(E) dE = 2g\pi V \left(\frac{2m}{h^2}\right)^{3/2} \int_0^{E_F} E^{3/2} dE$$
$$= \frac{4g\pi V}{5} \left(\frac{2m}{h^2}\right)^{3/2} E_{E_F}^{5/2} = \frac{2g\pi V}{5mh^3} p_F^5$$





Fig: Variation of chemical potential μ with *T*.

Strongly degenerate case for T > 0

• Chemical potential μ

$$\mu = E_F \left[1 - \left(\frac{T}{T_F} \right)^2 \frac{\pi^2}{12} \dots \right], \quad \text{where} \quad T_F = E_F / k_B = \mu / k_B \text{ is the Fermi temperature.}$$

Internal energy

$$U = 2g\pi V \left(\frac{2m}{h^2}\right)^{3/2} F_{5/2}(A) = 2g\pi V \left(\frac{2m}{h^2}\right)^{3/2} \frac{2}{5}\mu^{5/2} \left[1 + \left(\frac{kT}{\mu}\right)^2 \frac{5\pi^2}{8} + \dots\right]$$

 $\frac{U}{N} = \frac{3}{5} \frac{\mu^{5/2}}{E_F^{3/2}} \left[1 + \left(\frac{T}{T_F}\right)^2 \frac{5\pi^2}{8} + \dots \right]$ Using the value of μ , we can approximate

$$\frac{U}{N} = \frac{3}{5} E_F \left[1 - \left(\frac{T}{T_F}\right)^2 \frac{\pi^2}{12} \right]^{3/2} \left[1 + \left(\frac{T}{T_F}\right)^2 \frac{5\pi^2}{8} \right]$$
$$= \frac{3}{5} E_F \left[1 - \left(\frac{T}{T_F}\right)^2 \frac{5\pi^2}{24} \right] \left[1 + \left(\frac{T}{T_F}\right)^2 \frac{5\pi^2}{8} \right] = \frac{3}{5} E_F \left[1 + \left(\frac{T}{T_F}\right)^2 \frac{5\pi^2}{12} \right]$$

Thus, we have $U \approx \frac{3}{5} N E_F \left[1 + \frac{5\pi^2}{12} \left(\frac{T}{T_F} \right)^2 \right]$

• Hence, the specific heat at constant volume cv is

$$C_{V} = \left[\frac{\partial U}{\partial T}\right]_{V} = \frac{3}{5}NE_{F}\frac{5\pi^{2}}{6}\frac{T}{T_{F}^{2}} = k_{B}N\frac{\pi^{2}}{2}\left(\frac{T}{T_{F}^{2}}\right)$$

• Pressure of the ideal Fermi gas is

$$P = \frac{2}{3} \frac{U}{V} = \frac{3}{5} \frac{NE_F}{V} \left[1 + \frac{5\pi^2}{12} \left(\frac{T}{T_F} \right)^2 \right]$$

Example: Fermions of mass m are kept in two-dimensional box of area A at temperature T = 0

- (a) What is total number of particle if E_F is Fermi energy
- (b) What is the energy of the system if $E_{\rm F}$ is Fermi energy
- (c) Write expression of energy in term of E_F and N

For two dimensional systems density of state $g(E)dE = \pi A\left(\frac{2m}{h^2}\right)dE$ and distribution function

at temperature T = 0 for is given by f(E) = 1 if $E < E_F$

$$= 0$$
 if $E > E_F$

(a)
$$N = \int_0^{E_F} g(E) f(E) dE \implies N = \frac{\pi 2 m A E_F}{h^2}$$

(b)
$$E = \int_0^{E_F} Eg(E)f(E)dE$$
 $E = \frac{\pi 2mA}{h^2} \int_0^{E_F} EdE$, $E = \frac{\pi mAE_F^2}{h^2}$

(c)
$$E = \frac{NE_F}{2}$$

Example: (a) if Fermi gas is at temperature T > 0 what will $f(E_F)$

(b) At $E = E_F + x$, find the fraction of occupied levels

(c) At $E = E_F - x$, find fraction of unoccupied levels.

Solution: (a) It is also interesting to note that at T > 0, when $E = E_F$ we have

$$f(E) = \frac{1}{e^{E - E_F/kT} + 1} = \frac{1}{e^{0/kT} + 1} = \frac{1}{1+1} = \frac{1}{2}$$

(b) At T > 0, fraction of levels above E_F are occupied and a fraction of levels below E_F are vacant. The fraction of occupied levels at the energy E is

$$f(E) = \frac{1}{e^{(E-E_F)/kT} + 1}$$

At $E = E_F + x$, the fraction of occupied levels is $f(E_F + x) = \frac{1}{e^{x/kT} + 1}$ (c) The fraction of unoccupied levels at the energy E is

$$1 - f(E) = 1 - \frac{1}{e^{(E - E_F)/kT} + 1}$$

At $E = E_F - x$, the fraction of unoccupied levels is $1 - f(E_F - x) = \frac{1}{1 + e^{x/kT}}$